Optimal compensation of the Earth's magnetic field while changing the energy at FLASH

Martin Nuss¹

TU Graz, Austria

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¹Thanks to my supervisor Dr. Pedro Castro

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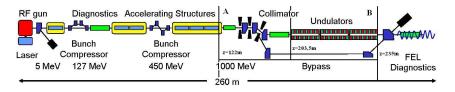
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FLASH

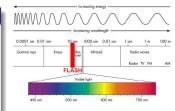
Tuning the Wavelength λ Changing the beam energy E Why does the trajectory change? SASE requires stable trajectory

FLASH



FLASH data

- Free Electron Laser: SASE
- length: $\approx 260m$
- electron beam energy: 0.3*GeV* to 1.0*GeV*
- synchrotron radiation: VUV and soft X-ray



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FLASH **Tuning the Wavelength** λ Changing the beam energy E Why does the trajectory change? SASE requires stable trajectory

Experimentalists work with different wavelengths



$$\lambda_{photons} = rac{\lambda_u}{2\gamma^2}(1+rac{\kappa}{2})$$

 γ ... relativistic γ factor K, λ_u ... undulator constants

FLASH has to tune energy of e^- beam to change λ

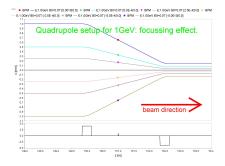
- $E = 0.3 GeV \Rightarrow \lambda = 70.0 nm$
- $E = 1.0 \text{GeV} \Rightarrow \lambda = 6.3 \text{nm}$

FLASH Tuning the Wavelength λ Changing the beam energy E Why does the trajectory change? SASE requires stable trajectory

Energy scaling procedure

Magnets (quadrupoles and corrector dipoles) are scaled by

$$\vec{I}_{new} = rac{E_{new}}{E_{initial}} \vec{I}_{initial}$$





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Compensating the Earth's magnetic field at FLASH

FLASH Tuning the Wavelength λ Changing the beam energy E Why does the trajectory change? SASE requires stable trajectory

Why does the trajectory change?



$$\vec{F}_L = q(\vec{E} + \vec{v} \times \vec{B})$$

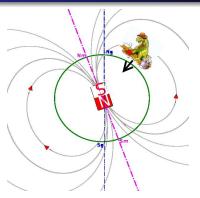
$$\vec{V} \quad \overrightarrow{\mathbf{B}} \quad \Rightarrow \quad \vec{F} = \vec{0}$$

Location of FLASH

- FLASH is situated almost in south-north direction
- $\vec{B}_{earth} || south north$
- ullet \Rightarrow no influence of Earth's magnetic field
- But what about vertical component of \vec{B}_{earth} ?

FLASH Tuning the Wavelength λ Changing the beam energy E Why does the trajectory change? SASE requires stable trajectory

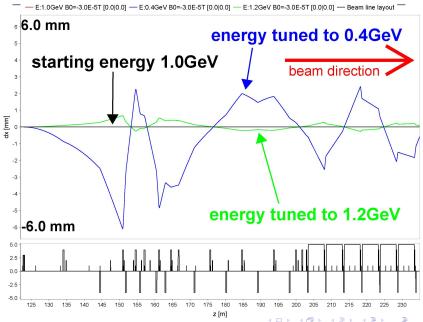
Effect of Earth's magnetic field



Hamburg (53°37'59"N/9°58'59"E)

- $\vec{B}_{earth,vertical} \approx -30 \mu T$
- $\bullet \Rightarrow \mathsf{vertical} \ \mathsf{component} \ \mathsf{affects} \ \mathsf{horizontal} \ \mathsf{trajectory}$

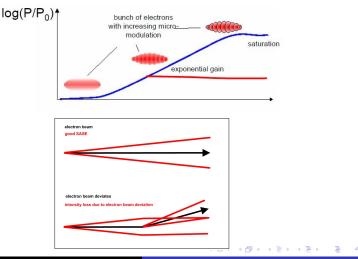
The trajectory while the beam energy changes



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Effect of the electron trajectory on SASE



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Procedure Methods



Procedure

- Simulate the effects of the Earth's magnetic field
- Find a compensation scheme

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Procedure Methods

Methods

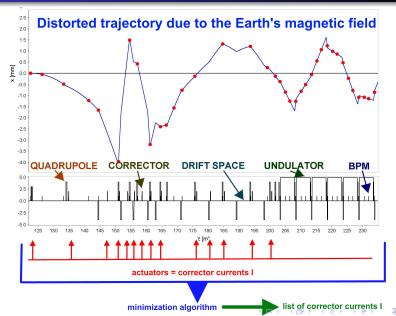
Tools

- Computer simulation
- Single particle motion in EM-fields
- Transport Matrix Formalism
- Numerical Optimization

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Calculation of the compensation for the Earth's magnetic field



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Procedure Methods

Applied corrector currents to compensate the Earth's magnetic field

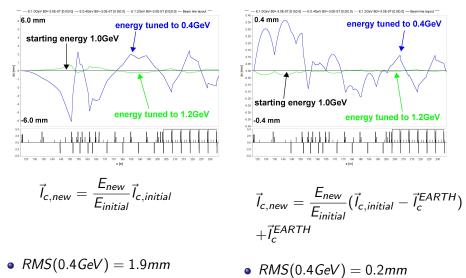
Corrector	Current	
H9ACC6	530.9	
H10ACC6	530.9	
H10ACC7	590.9	
H4TCOL	33.7	
H9TCOL	-4.4	
H2ECOL	57.7	
H4ECOL	25.6	
H6ECOL	89.3	
НЗМАТСН	5.5	
H6MATCH	18.7	
H5SUND2	25.7	
H4SUND3	<u>8</u> 1 [∢]	

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Compensating the Earth's magnetic field at FLASH

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Result, (Δx) - z is shown in the graph



• *RMS*(1.2*GeV*) = 0.21*mm*

• *RMS*(1.2*GeV*) = 0.01*mm*

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Result Conclusion



- An optimal compensation of the Earth's magnetic field has been found.
- A new scheme for changing the energy has been developed.
- The compensating currents \vec{l}_c^{EARTH} have been calculated.

Result Conclusion

Thank You!

Thank you for your attention!

If you have any questions feel free to ask them.

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