

## Fokker-Planck collision model and Monte Carlo method

The first part of the derivation follows closely the book *Collisional transport in magnetized plasmas* of Helander, Sigmar (2002), pp.22-24. Extensive information on the topic of kinetic theory and its various applications can be found in the book *Stochastic processes in physics and chemistry* of N.G. Van Kampen (2007).

We look at collisional effects to the velocity dependency of the distribution function  $f(v, t)$ . Since no particles are created or destroyed by collisions, after a sufficiently short time-step  $\Delta t$  the change in  $f$  at a fixed point  $v$  of velocity space can be written as

$$f(v, t + \Delta t) = \int d\Delta v f(v - \Delta v, t) F(v - \Delta v, \Delta v). \quad (1)$$

The function  $F(v - \Delta v, \Delta v)$  quantifies the probability with which a particle of velocity  $v - \Delta v$  will be scattered to have velocity  $v$ , and  $f(v - \Delta v, t)$  measures the population at the velocity space point from which scattering occurs. Integration corresponds to a summation over all possible distances  $\Delta v$  from  $v$ . The central assumption to obtain a Fokker-Planck collision operator is the sufficiently fast decay of  $F$  with  $\Delta v$ . Physically this means that most of the velocity changes by a single collision are small, which is a good approximation for Coulomb collisions in plasmas or for Brownian motion in chemistry (Kramers' equation). An expansion up to second order around  $v$  in the first argument  $v - \Delta v$  yields

$$f(v, t + \Delta t) \approx \int d\Delta v \left( f(v, t) F(v, \Delta v) - \Delta v \frac{\partial(f(v, t) F(v, \Delta v))}{\partial v} + \frac{\Delta v^2}{2} \frac{\partial^2}{\partial v^2} (f(v, t) F(v, \Delta v)) \right). \quad (2)$$

Since no particles can be lost, the integral over  $F(v, \Delta v)$  across the whole range of  $\Delta v$  has to fulfil the normalisation condition

$$\int \Delta v F(v, \Delta v) = 1, \quad (3)$$

which means that the first term can be integrated to yield just  $f(v, t)$ . Moving this term to the left-hand side and division by  $\Delta t$  yields the partial time derivative of  $f$  due to collisions in the limit  $\Delta t \rightarrow 0$  with

$$\begin{aligned} \left( \frac{\partial f}{\partial t} \right)_c &= \lim_{\Delta t \rightarrow 0} \frac{f(v, t + \Delta t) - f(v, t)}{\Delta t} \\ &\approx - \frac{\partial}{\partial v} \left( \underbrace{\frac{\langle \Delta v \rangle}{\Delta t}}_{A(v)} f(v, t) \right) + \frac{1}{2} \frac{\partial^2}{\partial v^2} \left( \underbrace{\frac{\langle \Delta v^2 \rangle}{\Delta t}}_{B(v)} f(v, t) \right), \end{aligned} \quad (4)$$

where we have defined

$$A(v) \equiv \frac{\langle \Delta v \rangle}{\Delta t}, \quad B(v) \equiv \frac{\langle \Delta v^2 \rangle}{\Delta t}$$

containing the integral over  $d\Delta v$  as moment averages for powers of  $\Delta v$  with weight  $F$ ,

$$\langle \Delta v^k \rangle = \int d\Delta v F(v, \Delta v) \Delta v^k. \quad (5)$$

If we furthermore assume  $A$  to be linear in  $v$  and  $B$  constant, we can write this collision term in the so-called Ornstein-Uhlenbeck form

$$\left( \frac{\partial f}{\partial t} \right)_c = \hat{L}_c f = \nu_c \frac{\partial}{\partial v} \left( v f + v_T^2 \frac{\partial f}{\partial v} \right), \quad (6)$$

where we have expressed Fokker-Planck coefficients  $A$  and  $B$  via collision frequency  $\nu_c$  and thermal velocity  $v_T$  to fulfil

$$A \equiv -v\nu_c, \quad (7)$$

$$B \equiv 2\nu_c v_T^2, \quad (8)$$

thus uniquely defining first and second moments  $\langle \Delta v \rangle$ ,  $\langle \Delta v^2 \rangle$  of  $\Delta v$  with respect to  $F$ . The choice of the name *thermal velocity* for  $v_T$  stems from the convergence of  $f$  towards a stationary thermalised state  $f_\infty$  given by a Gaussian in  $v$  (Boltzmann distribution in  $E = \frac{mv^2}{2}$  for 1D or Maxwellian for 3D) at  $t \rightarrow \infty$  with

$$f_\infty(v) = \frac{1}{\sqrt{2\pi v_T^2}} e^{-\frac{v^2}{2v_T^2}}. \quad (9)$$

It can be easily checked that  $\hat{L}_c f_\infty = 0$ , so a characterisation for the stationary  $f_\infty$  is being an eigenfunction of  $\hat{L}_c$  with eigenvalue 0.

To define a random process for a Monte Carlo method representing the Ornstein-Uhlenbeck operator, we require a random distribution of  $\Delta v$  whose first and second moment are given by Eqs. (7-8). Effectively this means to sample from a model for the distribution function  $F(\Delta v)$ . A straightforward choice is a normal (Gaussian) distribution

$$F(\Delta v) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{(\Delta v - \mu)^2}{2\sigma^2}} \quad (10)$$

with mean

$$\mu = \langle \Delta v \rangle = -\nu_c v \Delta t \quad (11)$$

and variance

$$\sigma^2 = \langle \Delta v^2 \rangle = 2\nu_c v_T^2 \Delta t. \quad (12)$$

In each time-step we draw a random sample from this distribution by

$$\Delta v = -\nu_c v \Delta t + \sqrt{2\nu_c v_T^2} \Theta \sqrt{\Delta t}, \quad (13)$$

where  $\Theta$  is a random number sampled from the standard normal distribution. As an alternative to using the normal distribution as a model for  $F$ , it would also be possible to use equally uniformly distributed random numbers centred around  $\mu$  and with the correct scaling of the variance.